

**PHYSICS EXAMINATION PROBLEMS
SOLUTIONS AND HINTS FOR STUDENT SELF-STUDY**

Module Code	PHY2201
Name of module	Statistical Physics
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1. (i) $\Delta S_{universe}^{reversible} = 0$. Melting ice: $\Delta S = mL_f/T = 1220 \text{ JK}^{-1}$. Heating water: $\Delta S = k_B \ln(T_2/T_1) = 1311 \text{ JK}^{-1}$.
- (ii) $\delta Q = 0 \Rightarrow C_V dT + \frac{A}{V^2} dV + PdV \Rightarrow C_V \frac{dT}{T} = -R \frac{dV}{V-B} \Rightarrow \ln[T^{C_V}(V-B)^R] = \text{const} \Rightarrow (RT)^{C_V}(V-B)^R = \text{const}$
 $\Rightarrow \left(P + \frac{A}{V^2}\right)^{C_V} (V-B)^{C_V+R} = \text{const} \Rightarrow \left(P + \frac{A}{V^2}\right) (V-B)^\gamma = \text{const}$, where $\gamma = \frac{C_V+R}{C_V}$.

2. See course notes. Isotropy implies $p(-u_x) = p(u_x)$.

See course notes. Note, that the number of states with the speed from u to $u + du$ is $2\pi u du$.

$$\varepsilon = mu^2/2 \Rightarrow u du = \frac{1}{m} d\varepsilon \Rightarrow p(u)du = 2 \frac{\alpha}{m} \exp(-2\alpha\varepsilon/m) d\varepsilon = \beta \exp(-\beta\varepsilon) d\varepsilon.$$

Boltzmann factor: $\exp(-\beta\varepsilon)$. $\langle \varepsilon \rangle = \int_0^\infty \varepsilon p(\varepsilon) d\varepsilon = \frac{1}{\beta} = 2 \frac{k_B T}{2} \Rightarrow \beta = \frac{1}{k_B T}$.

Density of states in 2D does not depend on energy. $p(\varepsilon < k_B T) = \int_0^{k_B T} p(\varepsilon) d\varepsilon = \int_0^1 e^{-t} dt = (e-1)/e$.

3. See course notes. $p_i = \exp\left(-\frac{\varepsilon_i}{k_B T}\right) / \sum_l \exp\left(-\frac{\varepsilon_l}{k_B T}\right)$. $Z = \sum_l \exp\left(-\frac{\varepsilon_l}{k_B T}\right)$ ensures $\sum_i p_i = 0$.

$$\bar{U} = \sum_r \varepsilon_r p_r = \frac{1}{Z} \sum_r \varepsilon_r \exp\left(-\frac{\varepsilon_r}{k_B T}\right) = -\frac{1}{Z} \frac{d}{d\beta} \sum_r \exp(-\beta\varepsilon_r) = -\frac{1}{Z} \frac{dZ}{d\beta} = -\frac{1}{Z} \frac{d \ln(Z)}{d\beta}.$$

$$Z = g_0 \exp\left(-\frac{\varepsilon_0}{k_B T}\right) + g_1 \exp\left(-\frac{\varepsilon_1}{k_B T}\right) + g_2 \exp\left(-\frac{\varepsilon_2}{k_B T}\right) = 1 + 3e^{-1} + 5e^{-2} \approx 2.78.$$

$$\bar{U} = \frac{1}{Z} \sum_r \varepsilon_r g_r \exp\left(-\frac{\varepsilon_r}{k_B T}\right) = k_B T (0 + 1 \times 3e^{-1} + 2 \times 5e^{-2}) / Z \approx 0.88 k_B T.$$

4. (i) $H = U + PV$. $dH = TdS + VdP \Rightarrow dH = TdS = \delta Q$ (heat flow) for $P = \text{const}$.
- (ii) See course notes: $F = U - TS$; extensive; $dF = -PdV - SdT \Rightarrow dF = -\delta W$ for $T = \text{const}$.

(iii) $\Delta S_{total} = 0 \Rightarrow C \left(\ln \frac{T_f}{T_A} + \ln \frac{T_f}{T_B} + \ln \frac{T_f}{T_C} + \ln \frac{T_f}{T_D} + \ln \frac{T_f}{T_E} \right) = C \ln \left[T_f^5 / (T_A T_B T_C T_D T_E) \right] = 0$
 $\Rightarrow T_f^5 / (T_A T_B T_C T_D T_E) = 1 \Rightarrow T_f = (T_A T_B T_C T_D T_E)^{1/5}$.

5. (i) $S = k_B \ln \Omega$; see course notes. $\Omega_{A+B} = \Omega_A \cdot \Omega_B \Rightarrow S_{A+B} = k_B \ln \Omega_A + k_B \ln \Omega_B = S_A + S_B$

(ii) a) $\Omega_0 = 1 \Rightarrow S = 0$; b) $\Omega_1 = N = 10 \Rightarrow S = k_B \ln N = k_B \ln 10 \approx 3.18 \times 10^{-23} \text{ JK}^{-1}$;

c) $S = k_B \ln \Omega_4 = k_B \ln \left(\frac{N!}{4!(N-4)!} \right) = k_B \ln \left(\frac{10!}{4!6!} \right) = k_B \ln(210) \approx 7.38 \times 10^{-23} \text{ JK}^{-1}$.

d) Most probable: $N_{heads} = N_{tails} = 5 \Rightarrow S = k_B \ln \Omega_5 = k_B \ln \left(\frac{10!}{5!5!} \right) = k_B \ln(252) \approx 7.63 \times 10^{-23} \text{ JK}^{-1}$.

- (iii) See course notes.

Classical limit: $\exp\left(\frac{E_i - E_F}{k_B T}\right) \gg 1$. Therefore, $f(E_i) \propto \exp\left(-\frac{E_i}{k_B T}\right)$ - Boltzmann's factor.